

Smarr Formula and Extended First Law for AdS Black Holes

Includes variations in Λ

Thermodynamic parameter
conjugate to Λ

$$dM = \frac{\kappa}{8\pi} dA + \frac{\Theta}{8\pi} d\Lambda$$

Kastor, Ray & Traschen, "Enthalpy and the
Mechanics of AdS Black Holes," arXiv:0904.2765

Kastor, Ray & Traschen, "Smarr Formula and
an Extended First Law for Lovelock Gravity,"
arXiv:1005.5053

Outline

- 1) Euler equation in classical thermodynamics
- 2) Obtaining Smarr formula with $\Lambda = 0$
- 3) $\Lambda \neq 0$ via "enhanced" Komar integral relations

Killing potential is
new ingredient

1) Euler equation in classical thermodynamics

Smarr formula
is GR analogue

$$dE = TdS - PdV \quad \text{1st Law}$$

can express

$$E = E(S, V) \quad \text{with} \quad T = \frac{\partial E}{\partial S}, \quad P = -\frac{\partial E}{\partial V}$$

E, S and V are all extensive variables

applying Euler's theorem for homogeneous functions....

$$f(\alpha^p x, \alpha^q y) = \alpha^r f(x, y) \quad \rightarrow \quad r f(x, y) = p \left(\frac{\partial f}{\partial x} \right) x + q \left(\frac{\partial f}{\partial y} \right) y$$

Gives Euler equation

$$E = TS - PV$$

Combine with 1st law to obtain
Gibbs-Duhem formula

$$0 = SdT - VdP$$

Relation between
intensive variables

2) Smarr formula with $\Lambda = 0$

$$dM = \frac{\kappa}{8\pi} dA + \Omega dJ \quad \text{1st law for stationary black holes}$$

Redo scaling argument in D dimensions ...

$$M \propto L^{D-3}, \quad A \propto L^{D-2}, \quad J \propto L^{D-2}$$

Applying Euler's theorem to
gives Smarr formula

$$M(A, J)$$

$$r f(x, y) = p \left(\frac{\partial f}{\partial x} \right) x + q \left(\frac{\partial f}{\partial y} \right) y$$

$$(D - 3)M = (D - 2) \frac{\kappa A}{8\pi} + (D - 2)\Omega J$$

2) Smarr formula with $\Lambda = 0$

Can also derive Smarr formula via Komar integral relation

ξ^a Killing vector

Let
$$B^{ab} = \left(\frac{D-2}{8\pi} \right) \nabla^a \xi^b$$

then
$$B^{ab} = B^{[ab]}$$

$$\nabla_a \nabla^a \xi^b = -R^b_c \xi^c$$

$$\nabla_a B^{ab} = 0 \quad \text{For vacuum spacetimes}$$

On a spatial slice extending from the horizon to infinity

$$I = \int_{\partial\Sigma} dS_{ab} B^{ab} = 0$$

For a static black hole

$$I_\infty = (D-3)M$$

$$I_H = (D-2) \frac{\kappa A}{8\pi}$$



Smarr formula

3) Smarr formula with $\Lambda \neq 0$

Focus on static black holes

Apply Euler's theorem to $M = M(A, \Lambda)$

scalings $M \propto L^{D-3}, \quad A \propto L^{D-2}, \quad \Lambda \propto L^{-2}$

defining $\frac{\Theta}{8\pi} \equiv \frac{\partial M}{\partial \Lambda}$

gives $(D-3)M = (D-2)\frac{\kappa A}{8\pi} - 2\frac{\Theta \Lambda}{8\pi}$

Form of Smarr formula

$$dM = \frac{\kappa}{8\pi} dA + \frac{\Theta}{8\pi} d\Lambda$$

Extended 1st law

New term accounts for change in mass when Λ is varied.

What is Θ ?

3) Smarr formula with $\Lambda \neq 0$

$$(D - 3)M = (D - 2)\frac{\kappa A}{8\pi} - 2\frac{\Theta\Lambda}{8\pi}$$

What is Θ ?

$$dM = \frac{\kappa}{8\pi}dA + \frac{\Theta}{8\pi}d\Lambda$$

Can compute explicitly for Schwarzschild-AdS spacetimes



Reasonable first step, but doesn't reveal geometric significance.

Caldarelli et. al. (1999)
Wang et. al. (2006)
Sekiwa (2006)
Wang (2006)
Larranaga Rubio (2007)
G. Cardoso & Grass (2008)
Urano et. al. (2009)

We give geometric derivations of....

Extended 1st law using Hamiltonian perturbation theory techniques

Sudarsky & Wald (1992)

Smarr formula via enhanced Komar integral relation

Easier

3) Smarr formula for AdS black holes via enhanced Komar integral relation

Killing potential

Introduce Killing potential

$$\nabla_a \xi^a = 0 \quad \longrightarrow \quad \xi^b = \nabla_a \omega^{ab}$$

Einstein equation

$$R_{ab} = \frac{2\Lambda}{(D-2)} g_{ab} \quad \longrightarrow \quad \nabla_a \nabla^a \xi^b = -\frac{2\Lambda}{D-2} \xi^b$$

New Komar integral relation

$$I = \frac{D-2}{8\pi} \int_{\partial\Sigma} \left(\nabla^a \xi^b + \frac{2\Lambda}{D-2} \omega^{ab} \right) = 0$$

$$0 = \frac{D-2}{8\pi} \left[\int_{\infty} dS_{ab} \left(\nabla^a \xi^b + \frac{2\Lambda}{D-2} \omega^{ab} \right) - \int_H dS_{ab} \left(\nabla^a \xi^b + \frac{2\Lambda}{D-2} \omega^{ab} \right) \right]$$

Both these terms diverge at infinity due to AdS asymptotics, but sum is finite.

Need to tease apart contributions to M and \mathfrak{V} .

$$0 = \frac{D-2}{8\pi} \left[\int_{\infty} dS_{ab} \left(\nabla^a \xi^b + \frac{2\Lambda}{D-2} \omega^{ab} \right) - \int_H dS_{ab} \left(\nabla^a \xi^b + \frac{2\Lambda}{D-2} \omega^{ab} \right) \right]$$

For pure AdS there is no horizon integral.

The two divergent terms cancel precisely.

Suggests adding and subtracting ω_{AdS}^{ab} ← Pure AdS Killing potential

$$0 = \frac{D-2}{8\pi} \left[\int_{\infty} dS_{ab} \left(\left[\nabla^a \xi^b + \frac{2\Lambda}{D-2} \omega_{AdS}^{ab} \right] + \frac{2\Lambda}{D-2} [\omega^{ab} - \omega_{AdS}^{ab}] \right) - \int_H dS_{ab} \left(\nabla^a \xi^b + \frac{2\Lambda}{D-2} \omega^{ab} \right) \right]$$

Gives $(D-3)M$ Contributes to Θ

Gives $(D-2) \frac{\kappa A}{8\pi}$ Contributes to Θ

Note that we've done a background subtraction and a background addition

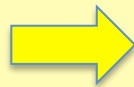
obtain $(D - 3)M = (D - 2)\frac{\kappa A}{8\pi} - 2\frac{\Theta\Lambda}{8\pi}$

with $\Theta = \int_H dS_{ab}\omega^{ab} - \int_\infty dS_{ab}(\omega^{ab} - \omega_{AdS}^{ab})$

Derivation of extended 1st
law yields same result

$$dM = \frac{\kappa}{8\pi}dA + \frac{\Theta}{8\pi}d\Lambda$$

What is Θ ?



“renormalized” difference of surface integrals of the Killing potential between horizon and infinity.

A few remarks ...

1. Θ is independent of ambiguities in choice of Killing potential

$$\xi^b = \nabla_a \omega^{ab}$$

2. For AdS-Schwarzschild one finds....

$$\Theta = -\frac{\Omega_{D-2} r_H^{D-1}}{D-1} = -V_H$$

Volume inside a sphere
of radius r_H in AdS

$$(D-3)M = (D-2)\frac{\kappa A}{8\pi} - 2\frac{\Theta\Lambda}{8\pi}$$

$$dM = \frac{\kappa}{8\pi} dA + \frac{\Theta}{8\pi} d\Lambda$$

$$\Theta = \int_H dS_{ab} \omega^{ab} - \int_\infty dS_{ab} (\omega^{ab} - \omega_{AdS}^{ab})$$

Agrees with previous
calculations of

$$\Theta \equiv 8\pi \frac{\partial M}{\partial \Lambda}$$

A few remarks ...

3. New term in first law relevant in a scenario with varying Λ .

- e.g. scenario of Brown & Teitelboim (1987) for neutralizing Λ through nucleation of pairs of branes charged under a D-form field strength.

- Perhaps there is some application for a thermodynamic ensemble at fixed ϑ , rather than fixed Λ ?

4. Can similarly derive Smarr formula and extended first law for asymptotically AdS black holes in Lovelock gravity.

- Perhaps interesting in the context of gauge-gravity duality with Gauss-Bonnet or Lovelock gravity in the bulk.

$$(D - 3)M = (D - 2)\frac{\kappa A}{8\pi} - 2\frac{\Theta\Lambda}{8\pi}$$

$$dM = \frac{\kappa}{8\pi}dA + \frac{\Theta}{8\pi}d\Lambda$$

$$\Theta = \int_H dS_{ab}\omega^{ab} - \int_{\infty} dS_{ab}(\omega^{ab} - \omega_{AdS}^{ab})$$

Includes variations in all the Lovelock coupling constants.

Bulk Lovelock couplings related to conformal anomaly terms on boundary.